

Darksky 2008

The 1st Continuous Measurements of Nightsky Brightness with a New Low Cost Luxmeter

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Motivation

- Quantitative statements
- 1st Continuous measurements
- Comparison between Jena – Obs. Tautenburg
- `Light weather`
- Long-time trends – disappearance of the Milky Way
- Integration in a meteorological station



Motivation

- Low costs (~ 100 Euro)
- Low maintenance
- Easy to operate
- Global network with 1000 stations
- Additional grid for `How many stars...`



Methods

- Conventional luxmeter
- Photometer (e.g. Zacharov, 1962)
- Simple digital camera (e.g. Schmidt, 2004)
- CCD-Camera (e.g. Cinzano et al., 2000)
- Imaging with fisheye lenses (e.g. Cristaldi et al., 2000)
- Anchor visual estimations (e.g. `How many stars...`)
- Solar cell (Kerschbaum/Posch)



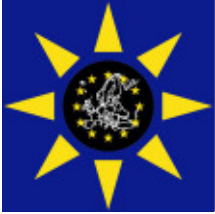
Methods - Disadvantages

- High costs
- Usage depends on weather
- Monitoring programs not always possible
- Complex data reduction and analysis
- All sky capability



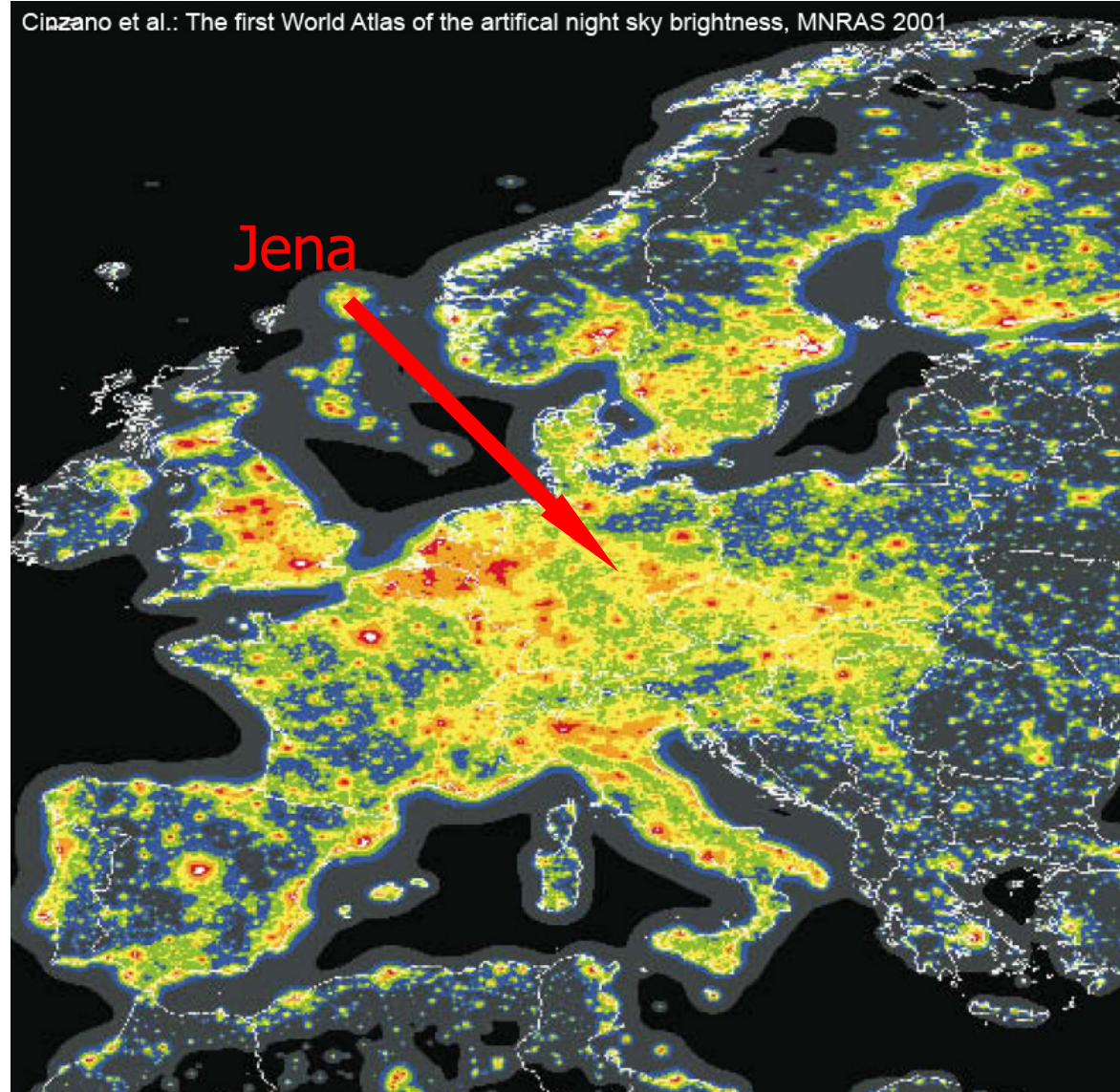
Our Low Cost Luxmeter

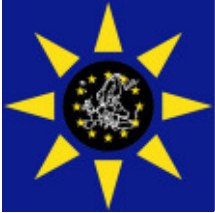
- Using a conventional solar cell as detector
- 1st prototype developed and built in summer 2005
- 1st light 26.04.2005
- Starting with monitoring program on 3rd Nov 2005



Measuring Site

Cinzano et al.: The first World Atlas of the artificial night sky brightness, MNRAS 2001





Measuring Site – Jena City



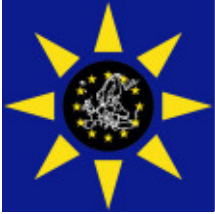
Operating for ~ 4 months

4m telescopic mast

transparent bowl for
weather protection

Pico-Ampere meter
located in temperature
stable staircase





Measuring Site – TLS Tautenburg

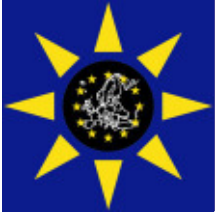




Solar Cell

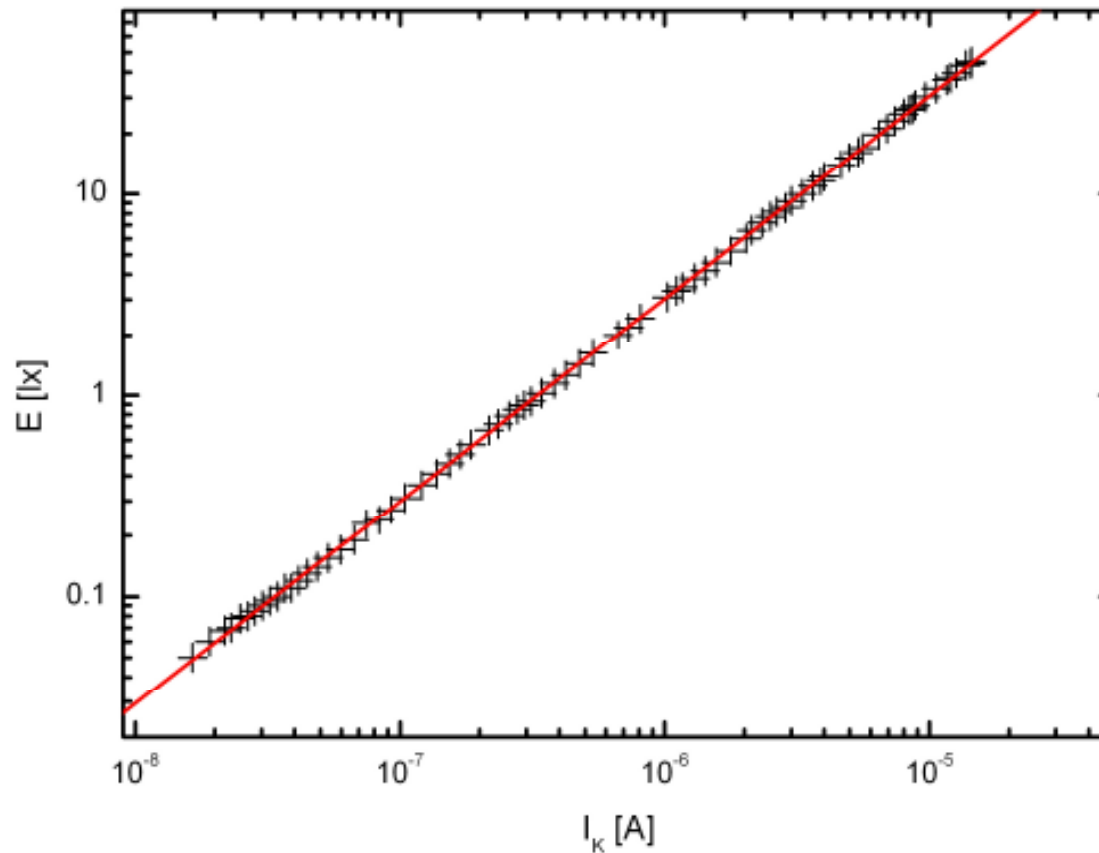
Advantages

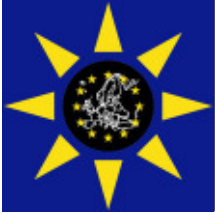
- Low costs (~ 100 Euro)
- Simple setup
- Continuous Measurements
- Weather independent
- Direction dependent measurements possible
- Could operate without PC and additional current (future)
- Mobility
- Measurements over 8 orders of magnitude (!)
- High resolution in time



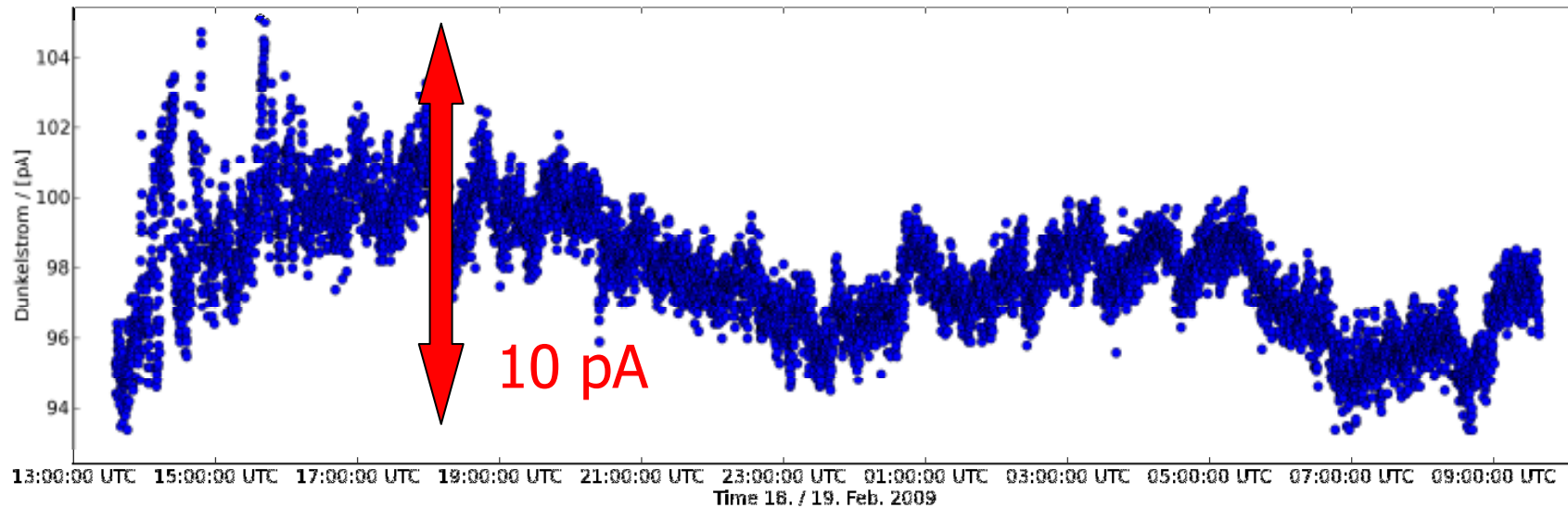
Calibration - Linearity

- For Calibration *VOLTCRAFT* (0.01 lx - 20.000 lx) was used
- In the future automatic calibration using the moon

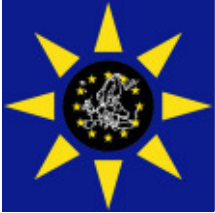




Dark Current

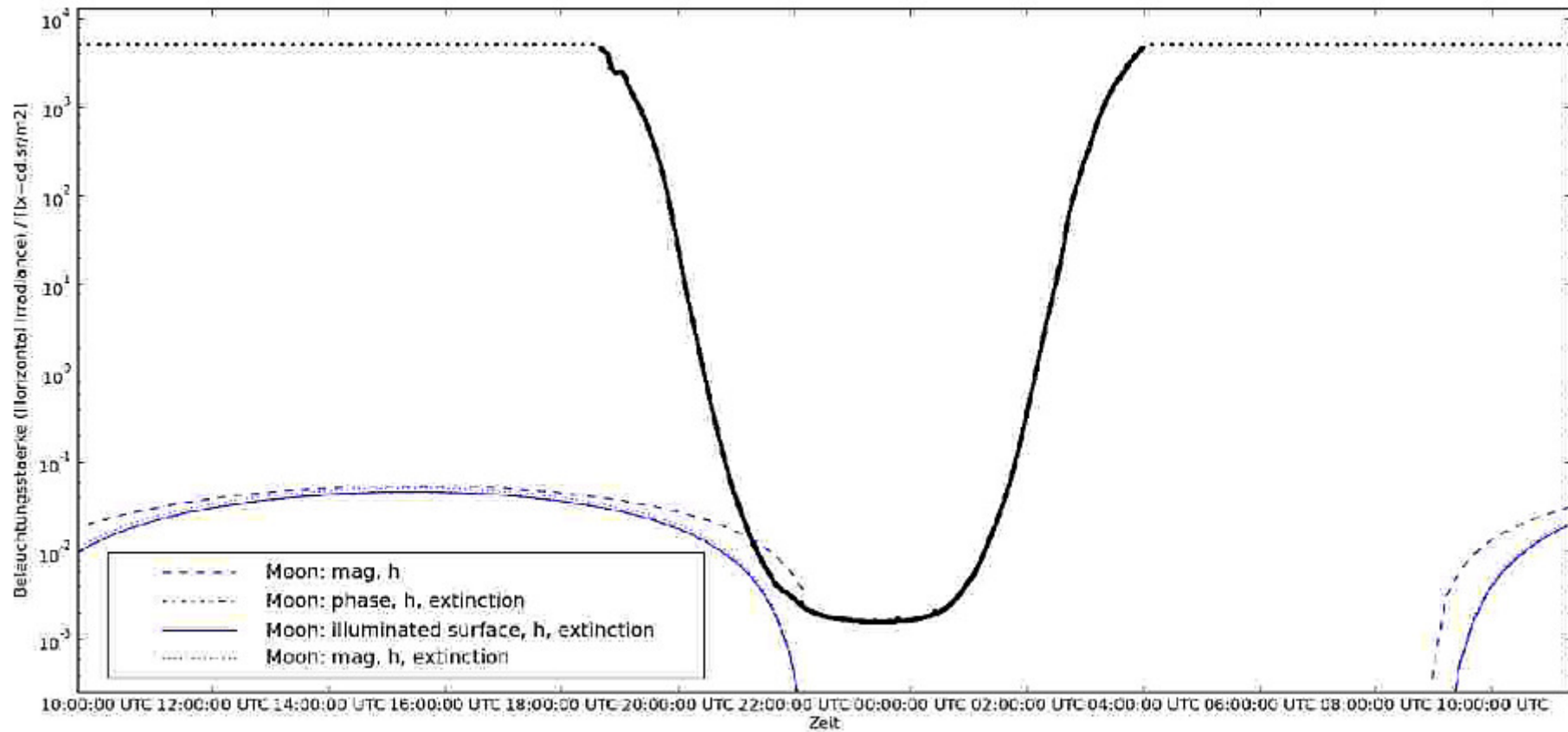


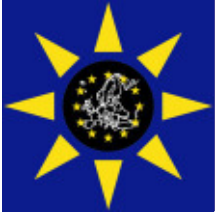
Lower detection limit $\sim 50\text{-}100 \mu\text{x}$
(natural night sky brightness $\sim 250 \mu\text{x}$)



First Results - Examples

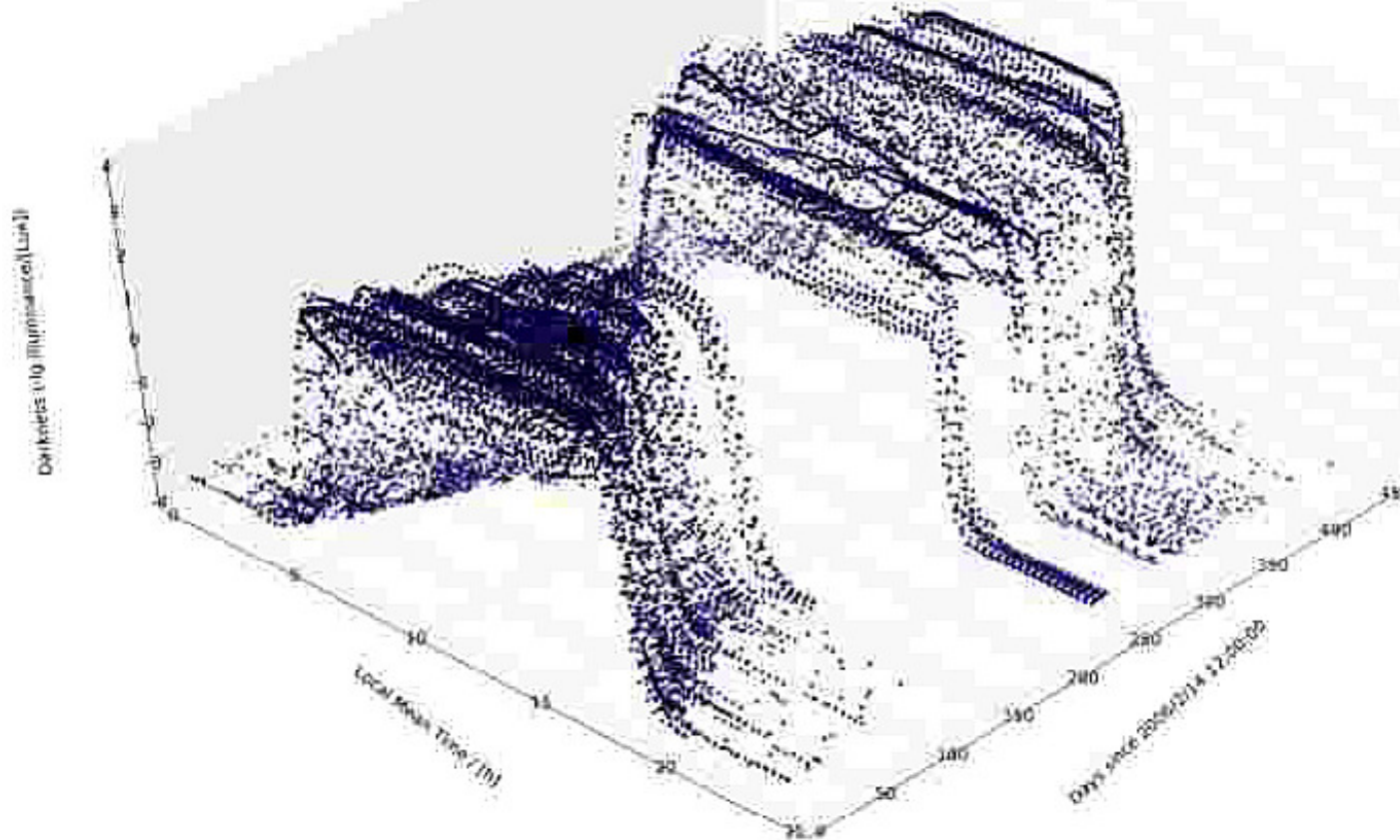
Darkest night in Tautenburg

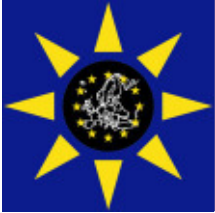




First Results - Examples

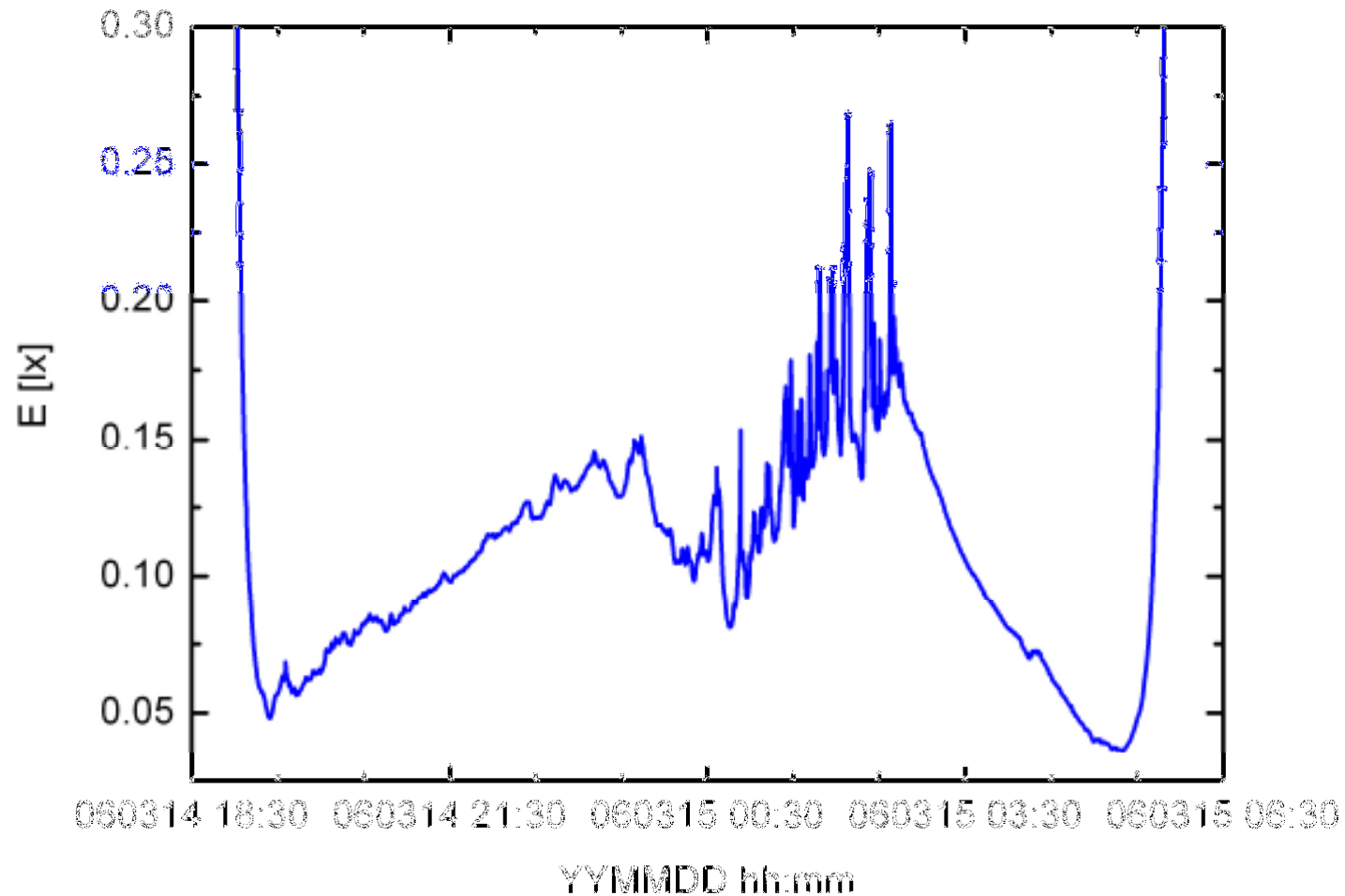
Change of night length
(Data from 2006, Tautenburg)

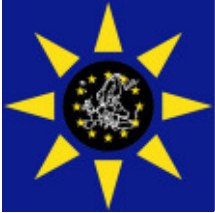




First Results - Examples

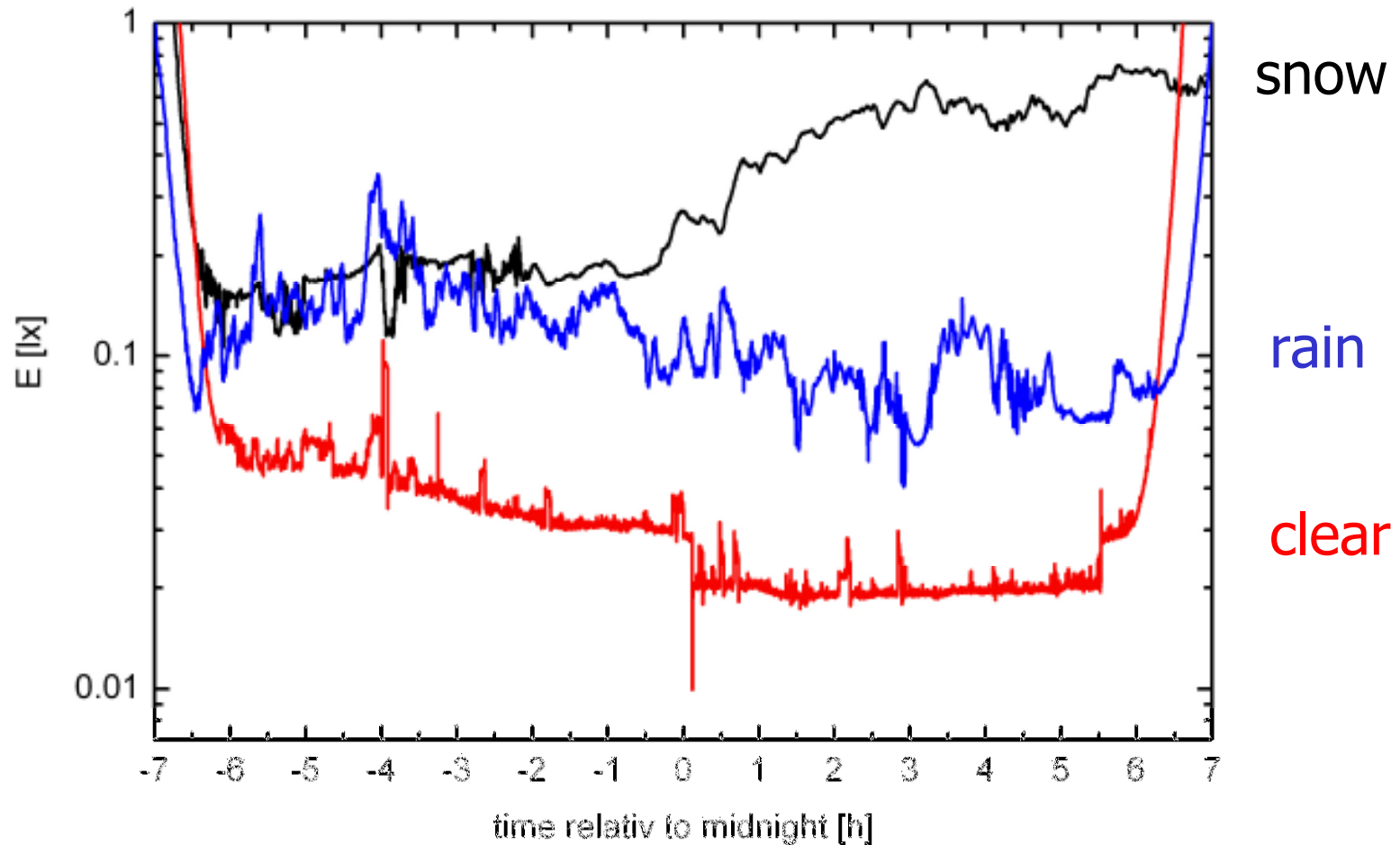
Night with full moon and clouds

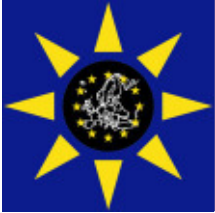




First Results - Examples

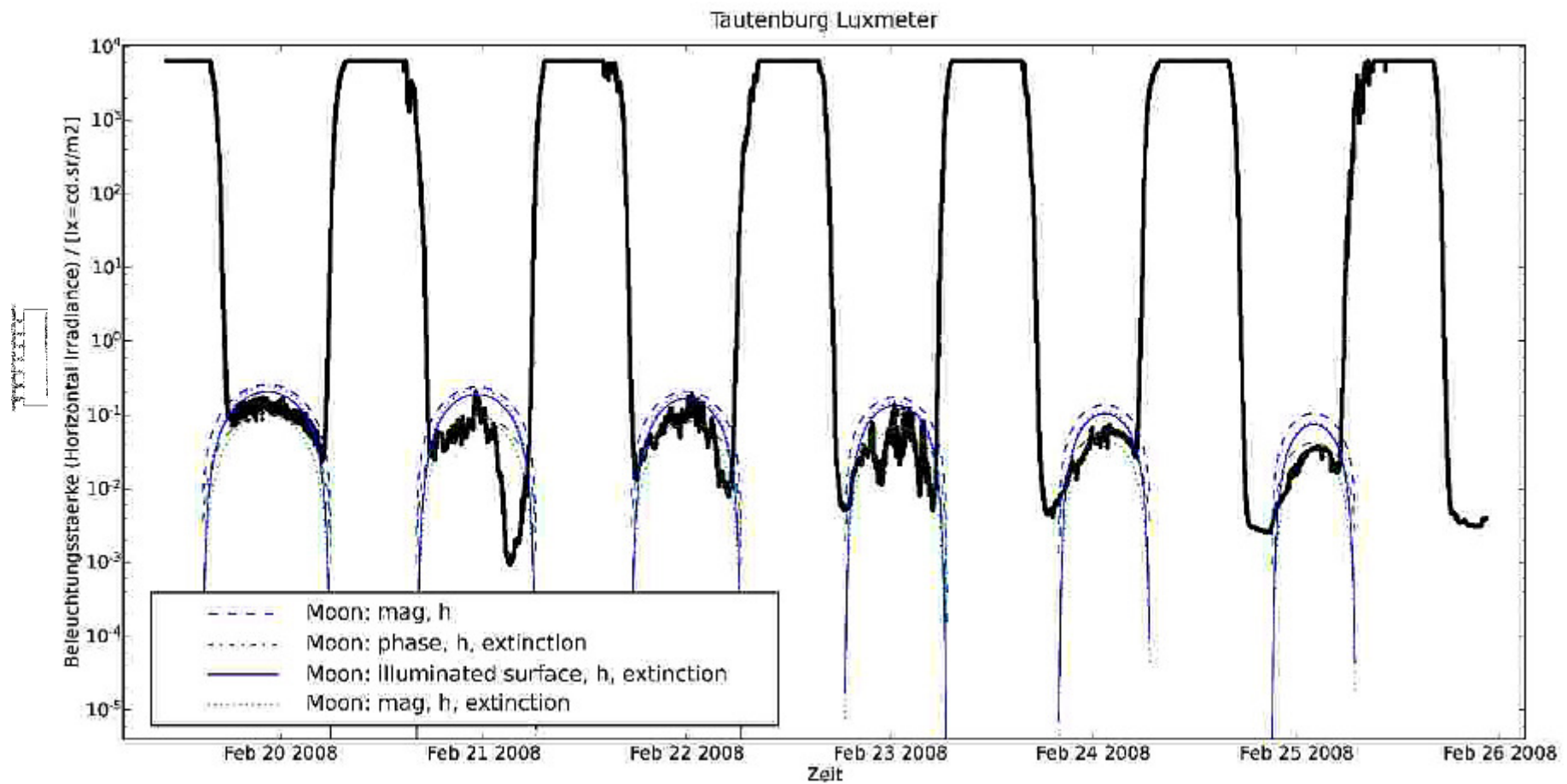
Influence of weather

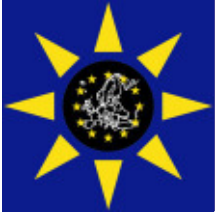




First Results - Examples

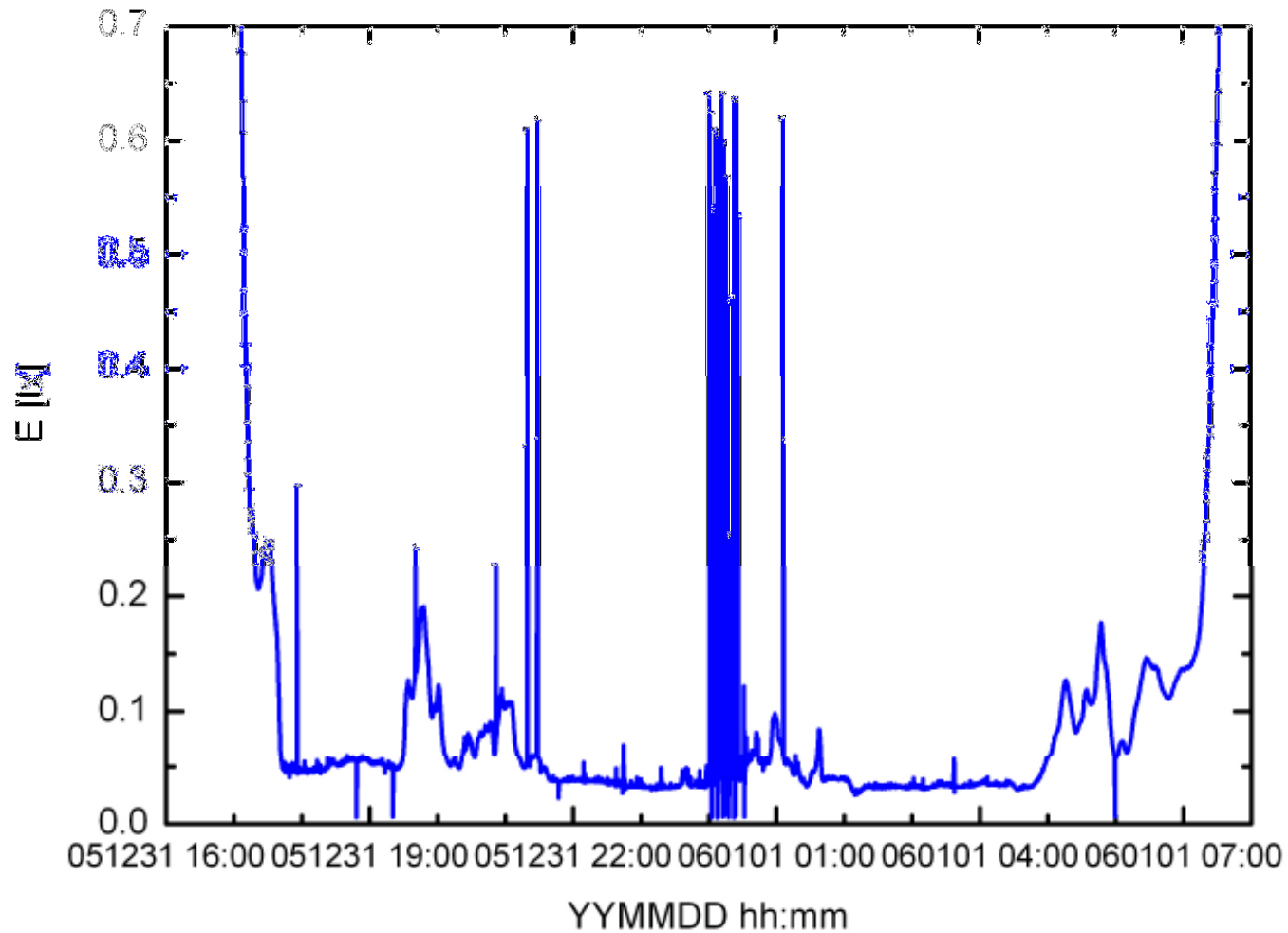
Modulation of illumination due to the moon

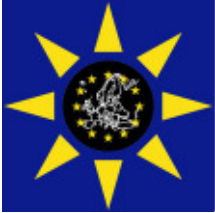




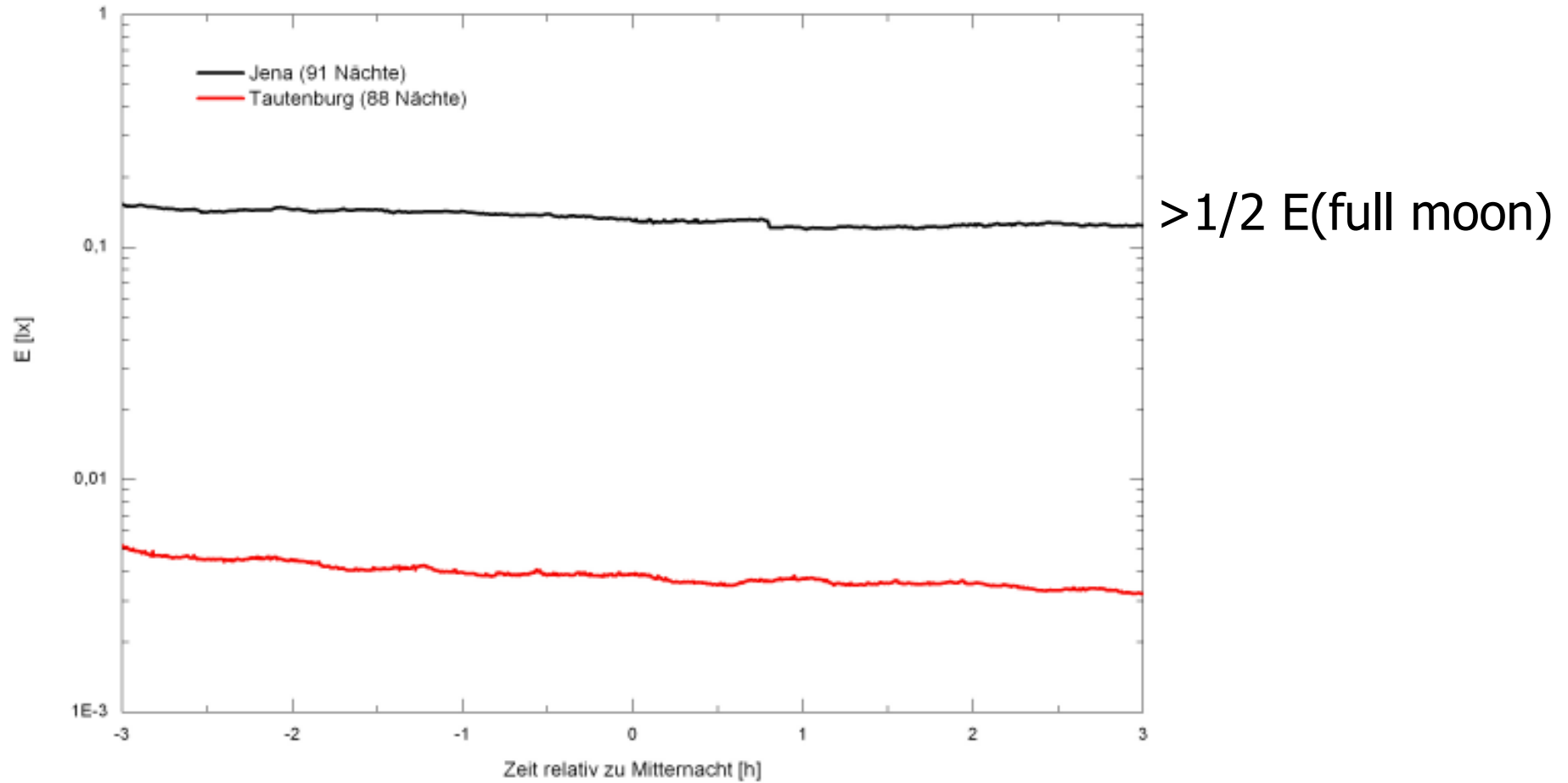
First Results - Examples

New Year's Eve 2005/2006

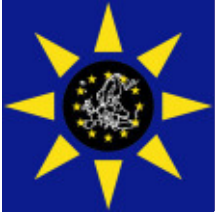




Jena - Tautenburg



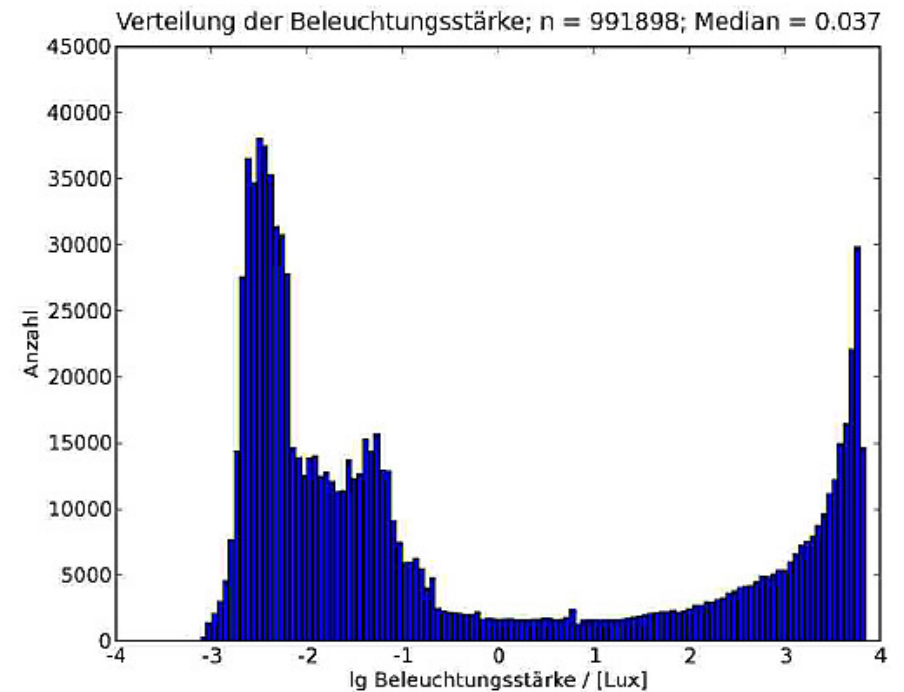
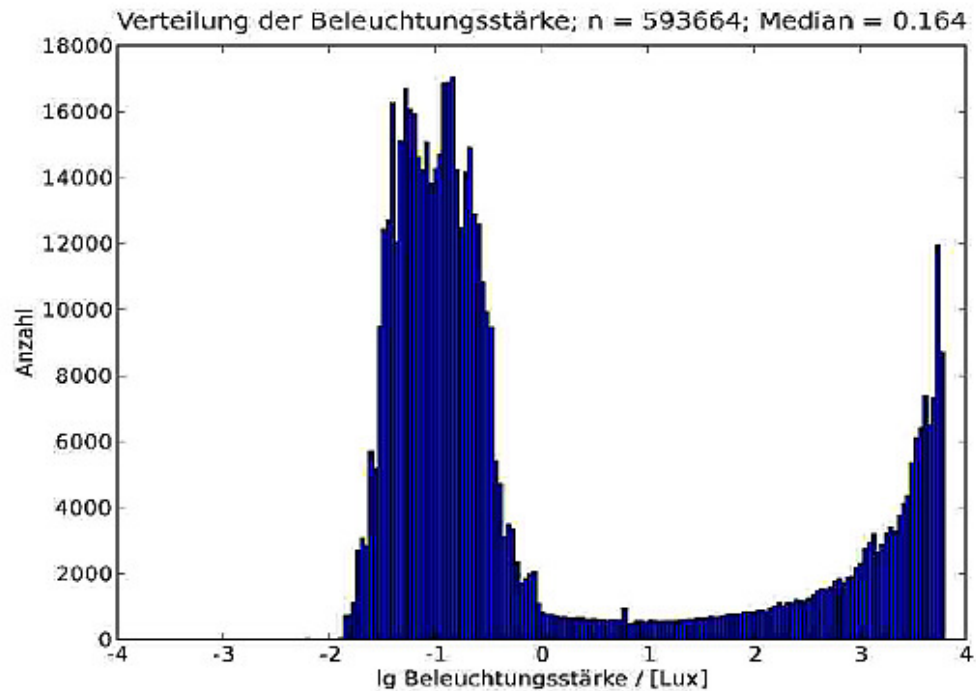
Jena is at least 40 times brighter than Tautenburg

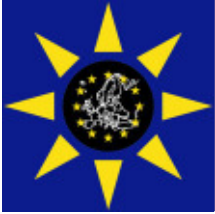


Jena - Tautenburg

Jena Nov 2005 – Jan 2006

Tautenburg 2006

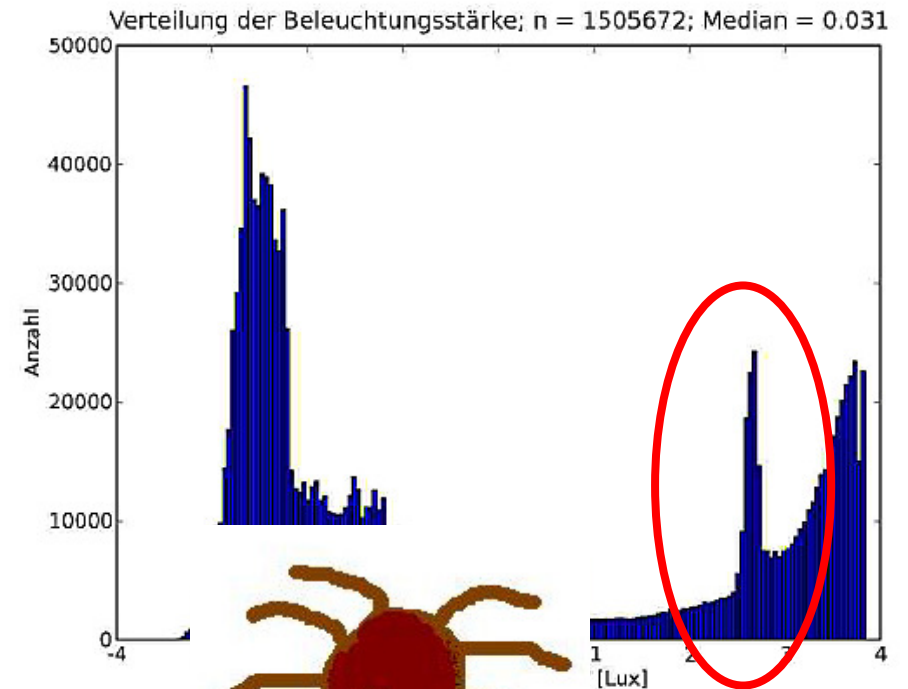
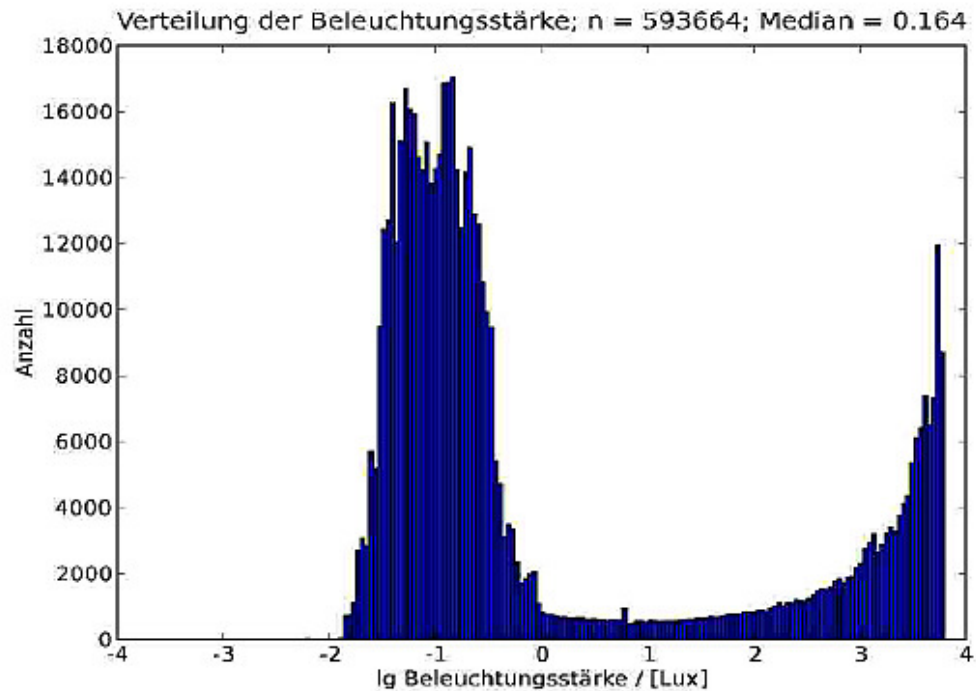




Jena - Tautenburg

Jena Nov 2005 – Jan 2006

Tautenburg 2007

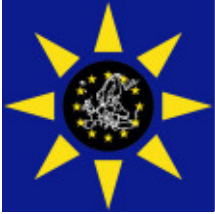




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Thank you for
your attention!





First Results - Examples

Night with full moon

